## **BRIEF COMMUNICATIONS ARISING**

# Causes of an AD 774-775 14C increase

ARISING FROM F. Miyake, K. Nagaya, K. Masuda & T. Nakamura Nature 486, 240-242 (2012)

Atmospheric  $^{14}\mathrm{C}$  production is a potential window into the energy of solar proton and other cosmic ray events. It was previously concluded that  $^{14}\mathrm{C}$  results from AD 774–775 would require solar events that were orders of magnitude greater than known past events  $^1$ . We find that the coronal mass ejection energy based on  $^{14}\mathrm{C}$  production is much smaller than claimed in ref. 1, but still substantially larger than the maximum historical Carrington Event of  $1859^{2-4}$ . Such an event would cause great damage to modern technology  $^{5.6}$ , and in view of recent confirmation of superflares on solar-type stars  $^{7.8}$ , this issue merits attention.

It was computed¹ that for a solar proton event (intercepted by the Earth) to account for the  $^{14}C$  data,  $\sim\!8\times10^{18}\,J$  of kinetic energy would be needed at the Earth, about 20 times that estimated³,⁴ for the historical Carrington Event of 1859. This was followed¹ by an estimate of  $2\times10^{28}\,J$  for the originating coronal mass ejection. As this is several orders of magnitude beyond the range⁵ of solar events, it was concluded¹ that solar causation of the  $^{14}C$  increase was not reasonable.

The fluence for this event at the Earth was calculated based on  $^{14}\text{C}$  production¹; we consider that the scaling applied to give the coronal mass ejection energy was incorrect. The applied geometrical factor was the ratio of the area of a sphere with radius the Earth's orbit to the cross-sectional area of the Earth. This would be appropriate if coronal mass ejections propagated isotropically. Instead, the opening angles are tyically  $^{10}$   $^{24}$  to  $^{20}$ , with smaller angles much more common. We use  $^{24}$ , corresponding to  $^{0.01}$  of the surface area of a sphere. The implied energy of the coronal mass ejection is now reduced to  $^{2}$   $^{26}$  J.

The correct scaling puts solar proton event energy at the lower end of observed solar-type star superflare energies<sup>7,8</sup>, which range from  $10^{26}$  J to  $10^{33}$  J, so there are solar-type stars with energy available to push coronal mass ejections well beyond the AD 774–775 event. Given the poor constraints on the rates of such large events at the Sun<sup>2,9</sup>, it would be wise to consider the possibility. A Carrington-level event would be disastrous for electromagnetic technology<sup>5,6</sup>, causing widespread damage to satellites and transformers linking the power grid. No assessment has been made of the technological effects of an event 20 times stronger.

Solar flare rates of occurrence follow a power-law in energy empirically based only for energies smaller than that of the Carrington Event. The probability of the implied event can be studied using the statistics of rare events<sup>11</sup>. Using the information that there was just one such event within  $\sim\!1,\!250\,\mathrm{yr}$ , the best estimate of the rate is  $8\times10^{-4}\,\mathrm{yr}^{-1}$ , with  $2\sigma$  (95.4%) confidence intervals from vanishingly small to  $3.2\times10^{-3}\,\mathrm{yr}^{-1}$ . Therefore the estimate of the probability of such an event within the next decade is 0.8%, which may be viewed as acceptably small unless the substantial technological consequences are considered.

Based on rate/energy scalings we have examined previously<sup>2</sup>, a short  $\gamma$ -ray burst could cause such effects from within  $\sim$ 1 kpc, but with an *a priori* probability of the order of  $10^{-4}$  over 1,250 yr.

Atmospheric ionization depletes ozone, increasing the solar UVB that reaches the ground<sup>2,3</sup>. We compute the ozone depletion<sup>12</sup> with corrected fluence and the results are shown in Fig. 1. This is not a mass extinction level event. The results are consistent with moderate biological effects: reduction of primary photosynthesis in the oceans and increased risk of erythaema and skin cancer, but no major mass-extinction level effects as implied earlier<sup>1</sup>. A newly published study<sup>13</sup> confirms our past computations of ozone depletion from a Carrington-level event, and suggests significant climate cooling

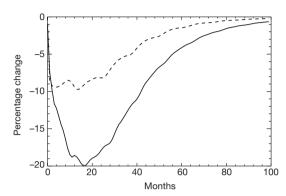


Figure 1 | Percentage change in globally averaged  $O_3$  column density. The change is between simulation runs with and without ionization input. Dashed line, with a short  $\gamma$ -ray burst; solid line, with a solar proton event.

might be a side effect, which would be enhanced for the event we describe here.

It is worth noting that the <sup>14</sup>C data could have been initiated by a series of solar proton events, each contributing a somewhat smaller amount. The <sup>14</sup>C production would add nearly linearly.

After these estimates were made, we noted a new study<sup>14</sup> of upper limits on energetic events at the Sun. An event of the energy resulting from our scaling lies just at their upper limits for an event that might appear every thousand years or so. Therefore a solar proton event appears to be a possible cause, which demands further exploration of a potential massive threat to modern civilization.

#### Adrian L. Melott<sup>1</sup> & Brian C. Thomas<sup>2</sup>

<sup>1</sup>Department of Physics and Astronomy, University of Kansas, Lawrence, Kansas 66045, USA.

e-mail: melott@ku.edu

<sup>2</sup>Department of Physics and Astronomy, Washburn University, Topeka, Kansas 66621, USA.

#### Received 9 August; accepted 4 October 2012.

- Miyake, F., Nagaya, K., Masuda, K. & Nakamura, T. A signature of cosmic-ray increase in AD 774–775 from tree rings in Japan. *Nature* 486, 240–242 (2012).
- Melott, A. L. & Thomas, B. C. Astrophysical ionizing radiation and the Earth: a brief review and census of intermittent intense sources. Astrobiology 11, 343–361 (2011).
- Clauer, C. R. & Siscoe, G. (eds) The great historical geomagnetic storm of 1859: a modern look. Adv. Space Res. 38, 115–388 (2006).
- Thomas, B. C., Jackman, C. H. & Melott, A. L. Modeling atmospheric effects of the September 1859 solar flare. Geophys. Res. Lett. 34, L06810, http://dx.doi.org/ abs/10.1029/2006GL029174 (2007).
- National Research Council Space Studies Board. Severe Space Weather Events Understanding Societal and Economic Impacts (National Academies Press, 2008); available at http://www.nap.edu/openbook.php?record\_id=12507&page=R1.
- Hapgood, M. Astrophysics: prepare for the coming space weather storm. Nature 484, 311–313 (2012).
- Schaefer, B. E., King, J. R. & Deliyannis, C. P. Superflares on ordinary solar-type stars. Astrophys. J. 529, 1026–1030 (2000).
- 8. Maehara, H, et al. Superflares on solar-type stars. Nature 485, 478–481 (2012).
- Reedy, R. C. in Solar Drivers of Interplanetary and Terrestrial Disturbances: Proceedings of the 16th International Workshop (eds Balasubramaniam, K. S., Keil, S. L. & Smartt, R. N.) 429–436 (ASP Conf. Ser. Vol. CS-94, Astronomical Society of the Pacific, 1996).
- Schrijver, C. J. Eruptions from solar ephemeral regions as an extension of the size distribution of coronal mass ejections. Astrophys. J. 710, 1480–1485 (2010).

### **BRIEF COMMUNICATIONS ARISING**

- Love, J. J. Credible occurrence probabilities for extreme geophysical events: earthquakes, volcanic eruptions, magnetic storms. *Geophys. Res. Lett.* 39, L10301, http://dx.doi.org/10.1029/2012GL051431 (2012).
- Ejzak, L. M., Melott, A. L., Medvedev, M. V. & Thomas, B. C. Terrestrial consequences of spectral and temporal variability in ionizing photon events. *Astrophys. J.* 654, 373–384 (2007).
- Calisto, M., Verronen, P. T., Rozanov, E. & Peter, T. Influence of a Carrington-like event on the atmospheric chemistry, temperature and dynamics. *Atmos. Chem. Phys.* 12, 8679–8686 (2012).
- Schrijver, C. J. et al. Estimating the frequency of extremely energetic solar events, based on solar, stellar, lunar, and terrestrial records. J. Geophys. Res. 117, A08103, http://dx.doi.org/10.1029/2012JA017706 (2012).

**Author Contributions** A.L.M. planned and wrote the paper with the assistance of B.C.T. B.C.T. performed the atmospheric computations and made the plot.

Competing Financial Interests Declared none.

doi:10.1038/nature11695